

Regularity of linear waves at the Cauchy horizon of black hole spacetimes

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joint with András Vasy

Luminy
April 29, 2016

Cauchy horizon of charged black holes

(subextremal) Reissner-Nordström-de Sitter spacetime

- ▶ solution of Einstein-Maxwell system,
- ▶ black hole mass $M > 0$, charge $Q > 0$,

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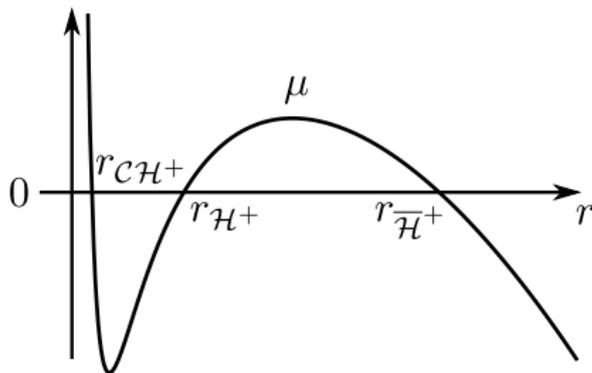
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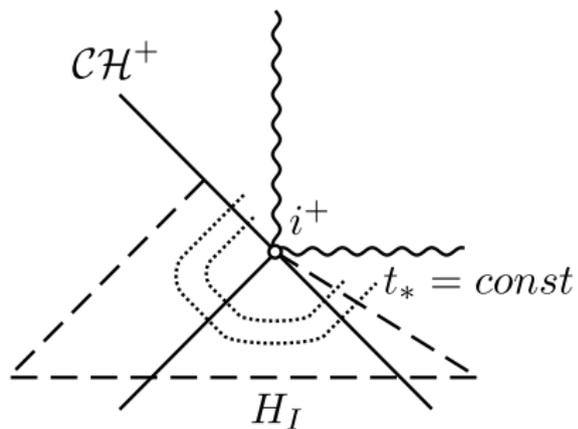
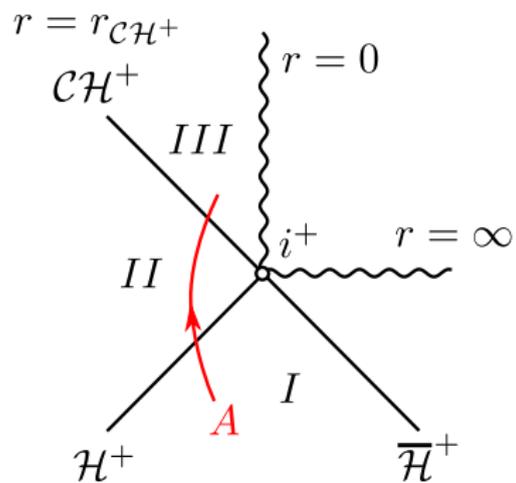
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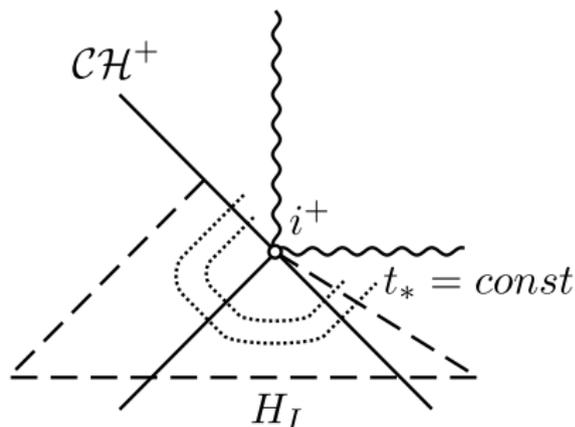
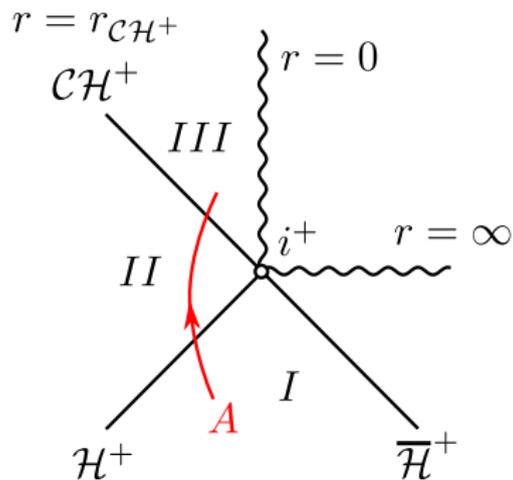
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- ▶ cosmological constant $\Lambda > 0$,
- ▶ topology: $\mathbb{R}_{t_*} \times (0, \infty)_r \times \mathbb{S}^2$,
- ▶ metric: $g = \mu(r) dt^2 - \mu(r)^{-1} dr^2 - r^2 d\sigma^2$; $t_* = t - F(r)$.



Penrose diagram:



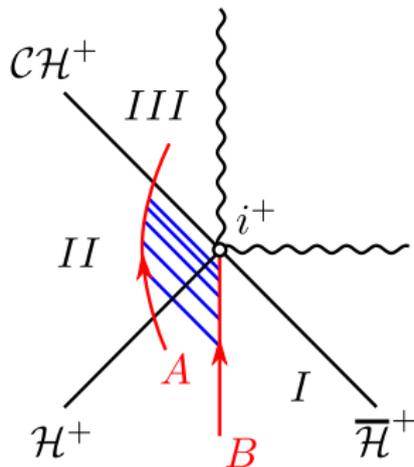
Penrose diagram:



Cauchy horizon \mathcal{CH}^+ : boundary of domain of uniqueness of solution u to wave equation $\square_g u = 0$ with Cauchy data on H_I

Blue-shift effect and strong cosmic censorship

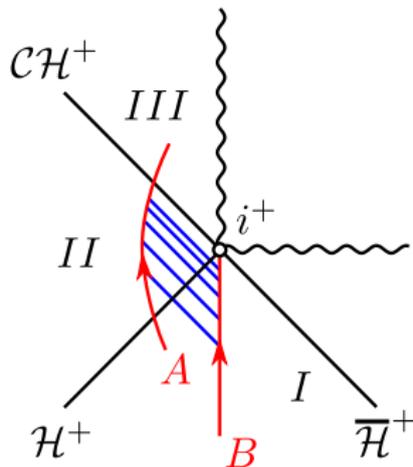
(Simpson–Penrose '73.)



Observer **A** crosses \mathcal{CH}^+ in finite time.

Blue-shift effect and strong cosmic censorship

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Observer A crosses \mathcal{CH}^+ in finite time.

Observer B lives forever.

Conjecture (Penrose's Strong Cosmic Censorship)

The maximal globally hyperbolic development of generic initial data for Einstein's field equations is inextendible as a suitably regular Lorentzian manifold.

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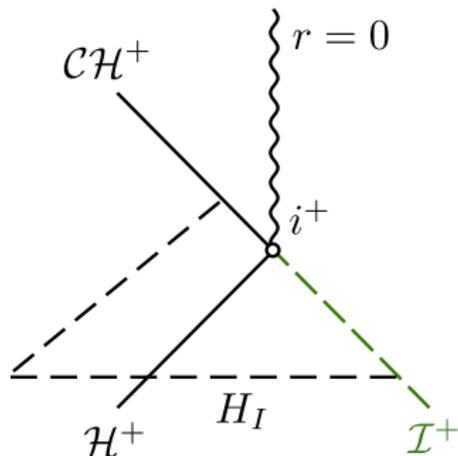
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Related work:

- ▶ Christodoulou (. . . , '99, '08),
- ▶ Dafermos ('03, '05, '13),
- ▶ ongoing work by Dafermos–Luk, Luk–Oh.

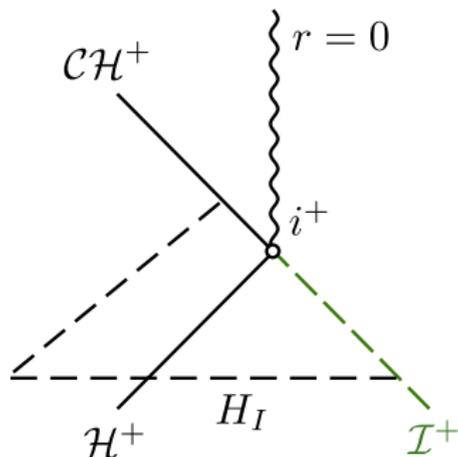
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Cauchy problem for $\square_g u = 0$ on
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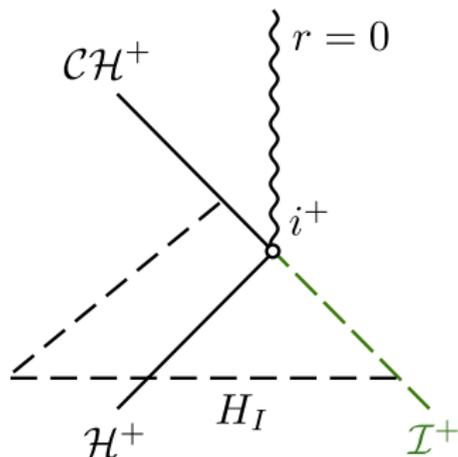


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For C^∞ initial data, u remains uniformly bounded near \mathcal{CH}^+ .

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Theorem (Luk–Oh, '15)

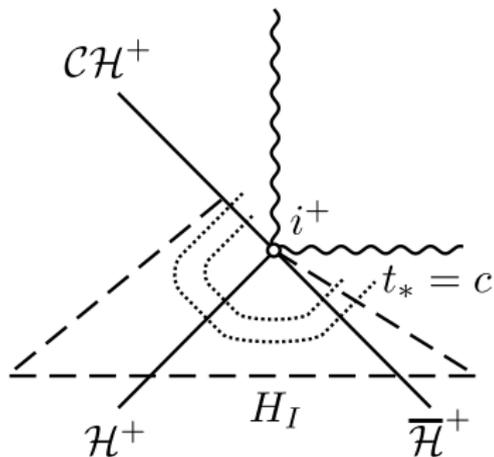
For generic C^∞ initial data, u is **not** in H_{loc}^1 near any point of \mathcal{CH}^+ .

Theorem (H.-Vasy, '15)

For C^∞ initial data on
Reissner–Nordström–de Sitter, u
solving $\square_g u = 0$ has a partial
asymptotic expansion near i^+ ,

$$u = u_0 + u', \quad u_0 \in \mathbb{C}, \quad |u'(t_*)| \lesssim e^{-\alpha t_*},$$

where $\alpha > 0$ depends only on the
spacetime.

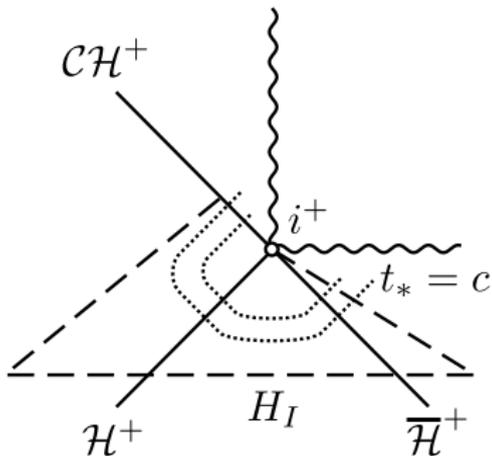


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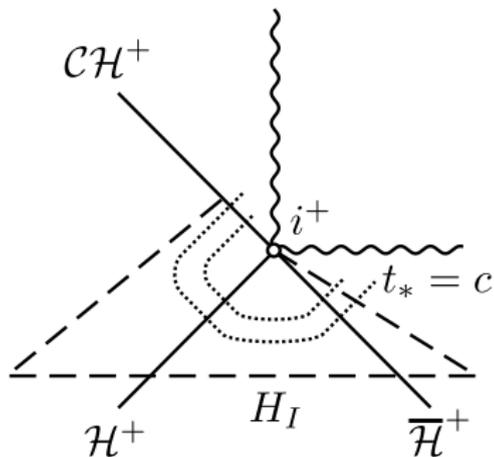
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$\kappa > 0$: surface gravity of the Cauchy
horizon



Previous work

Microlocal analysis/scattering theory approach:

- ▶ Melrose ('93)
- ▶ Sá Barreto–Zworski ('97), Bony–Häfner ('08)
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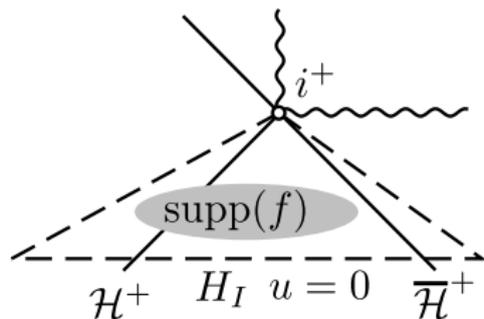
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Energy estimates:

- ▶ Dafermos–Shlapentokh–Rothman ('15)
- ▶ Luk–Sbierski ('15)

Analysis near the exterior region

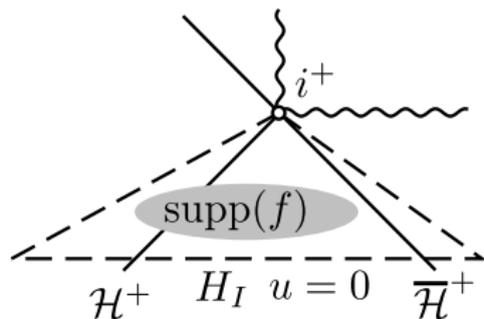


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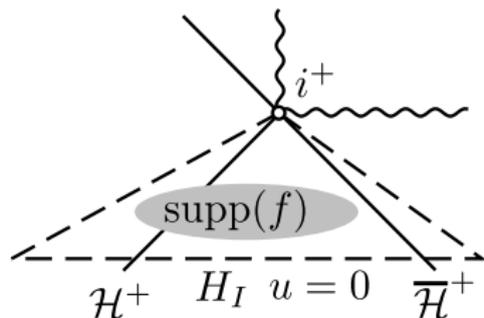


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u is C^∞ . Quantitative bounds as $t_* \rightarrow \infty$?

Microlocal red-shift effect

$\square_g u = f \in \mathcal{C}_c^\infty$, u smooth.

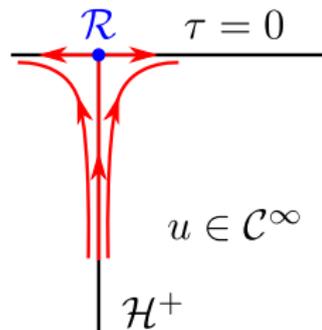
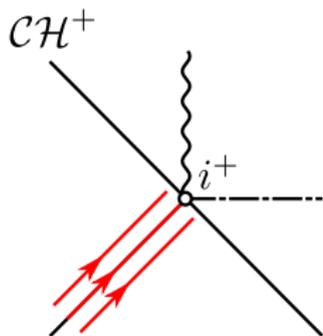
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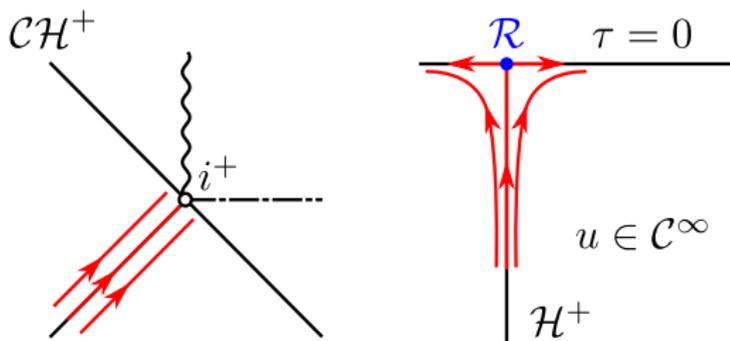
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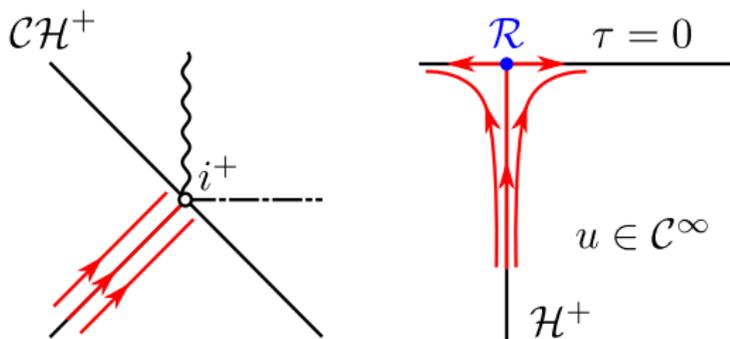
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Resonance expansions

'Spectral' family $\widehat{\square}_g(\sigma) = e^{it_*\sigma} \square_g e^{-it_*\sigma}$.

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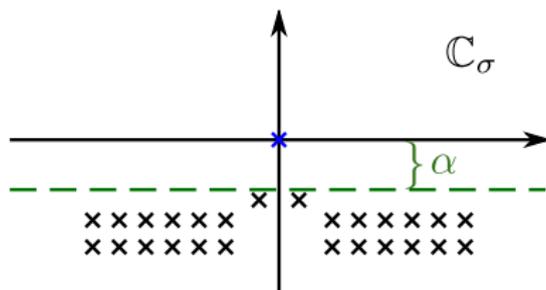
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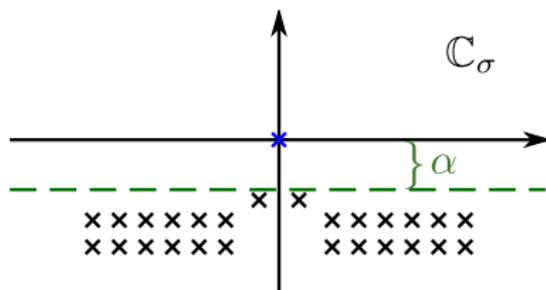
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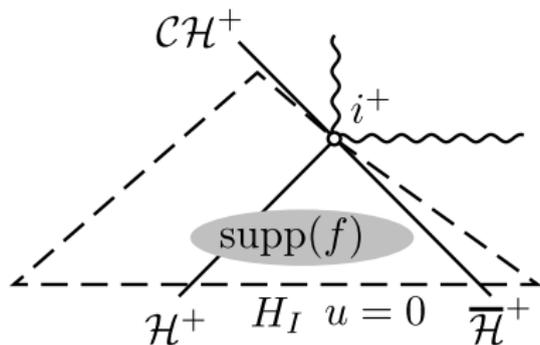
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Obtain:

$$u = u_0 + u', \quad u_0 \in \mathbb{C}, \quad u' \in e^{-\alpha t_*} H^\infty.$$

This gives asymptotics and decay in $r \geq r_{\mathcal{CH}^+} + \epsilon$, $\epsilon > 0$.



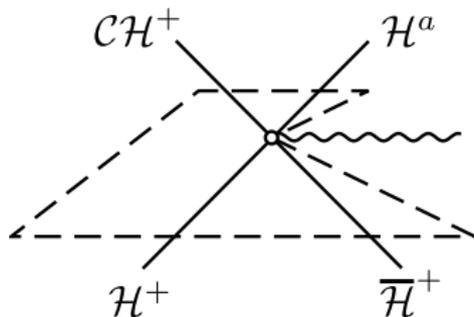
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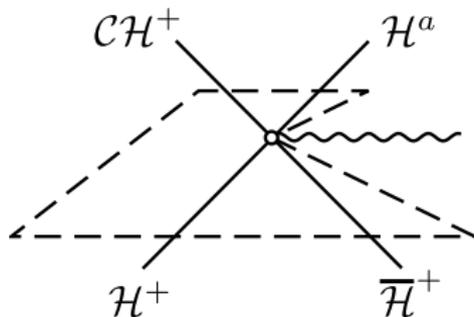
Modify spacetime beyond \mathcal{CH}^+ : Add **artificial exterior region**.



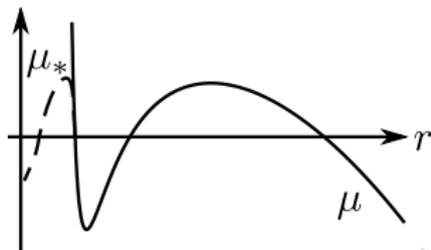
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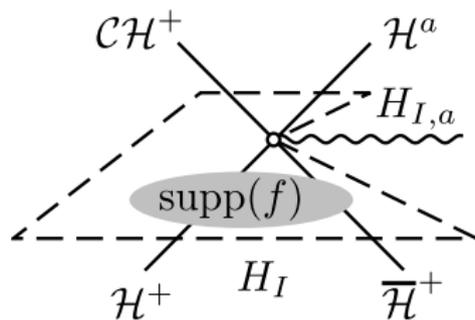


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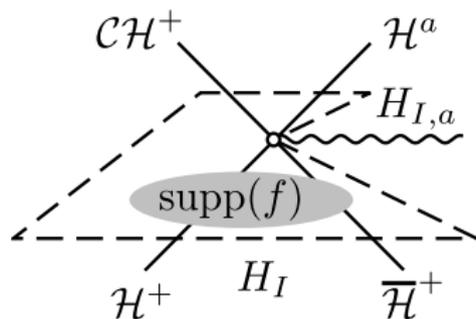
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Study forcing problem $\square_{\tilde{g}} u = f$, with $u = 0$ near $H_I \cup H_{I,a}$.



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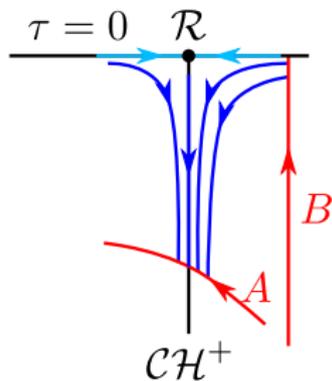
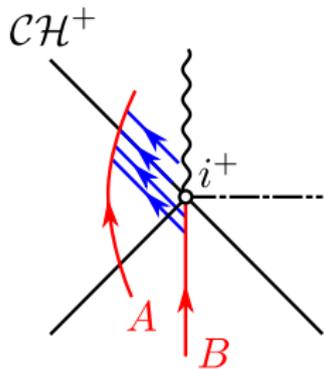
(Add **complex absorbing potential** $Q \in \Psi_b^2$ beyond \mathcal{CH}^+ to hide additional trapping and \mathcal{H}^a . Study $\square_{\tilde{g}} - iQ$.)

Microlocal blue-shift effect

□ $\tilde{g}u = f \in \mathcal{C}_c^\infty$. Work near \mathcal{CH}^+ . Recall $\tau = e^{-t_*}$.

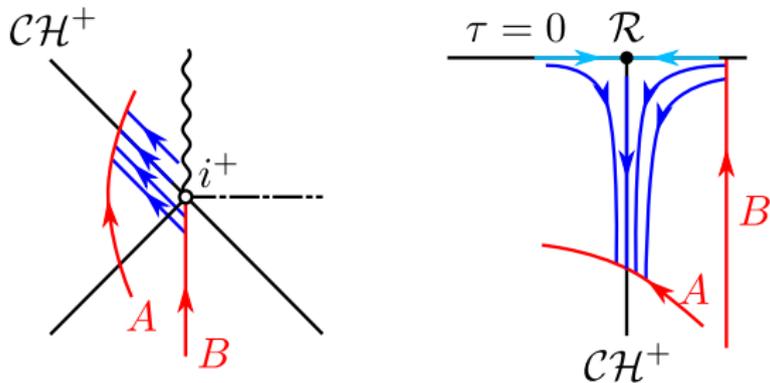
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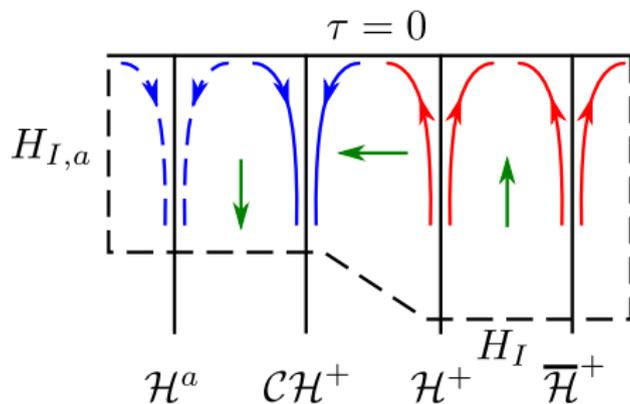
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Microlocal analysis of the extended problem



Green arrows: Future directed timelike vectors.

Obtain

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(s is a variable order function, and $\ell < 0$.)

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Get solvability of extended problem $\square_{\tilde{g}} u = f$, and **control of u near \mathcal{CH}^+** .

Resonance expansion

Solution u of $\square_{\tilde{g}} u = f$ has partial expansion

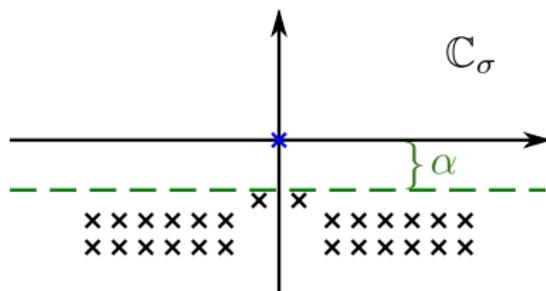
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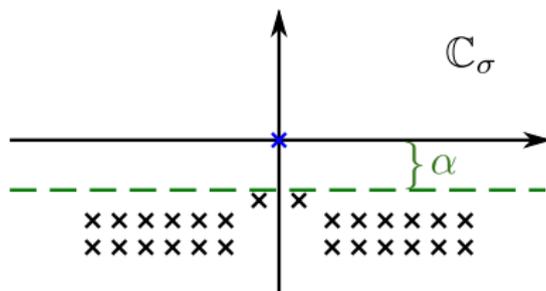


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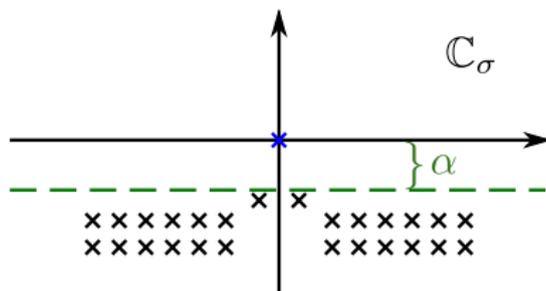
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$H^{1/2+0}(\mathbb{R}_r) \hookrightarrow L^\infty(\mathbb{R}_r)$ yields $|u'(t_*)| \lesssim e^{-\alpha t_*}$.

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Only potential issue for large a : resonances in $\text{Im } \sigma \geq 0$ ('mode stability,' see Whiting '89, Shlapentokh-Rothman '14 for Kerr)

Outlook

Shallow resonances. Mode solution $\square_g(e^{-i\sigma t_*} v(x)) = 0$ has $v \in H^{1/2 - \text{Im } \sigma / \kappa - 0}$ at \mathcal{CH}^+ ; could be C^∞ in principle. Study **location and regularity properties** of shallow resonances.

Outlook

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- Nonlinear problems.** Einstein's field equations. Work in progress by Luk–Rodnianski, Dafermos–Luk, Luk–Oh.